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In this paper we address the problem of the arrow of time from a cosmological point of view, rejecting the traditional entropic approach that defines the future direction of time as the direction of the entropy increase: from our perspective, the arrow of time has a global origin and it is an intrinsic, geometrical feature of space-time. Time orientability and existence of a cosmic time are necessary conditions for defining an arrow of time, which is manifested globally as the time-asymmetry of the universe as a whole, and locally as a time-asymmetric energy flux. We also consider arrows of time of different origins (quantum, electromagnetic, thermodynamic, etc.) showing that they can be non-conventionally defined only if the geometrical arrow is previously defined.

I. INTRODUCTION

In our previous papers on time-asymmetry ([1], [2], [3], [4], [5], [6], [7], [8]), the main features of this problem were considered (e.g. the thermodynamic arrow of time in early universe [7] or the arrow of time in some cosmological models [8]). In this paper we will try to present a comprehensive and updated formulation of the subject from a cosmological point of view.

As we have said many times before, the problem of time-asymmetry, also known as the problem of the arrow of time, can be formulated by means of the following question: *How an evident time-asymmetry is possible if the laws of physics are time-reversal invariant?* In fact, the laws of physics are invariant under the transformation $t \rightarrow -t$ ¹. Nevertheless, we have the psychological feeling that past is different than future; moreover, there are clear time-asymmetric phenomena, being the natural tendency from non-equilibrium to equilibrium the most conspicuous example. Astonishing enough, the solution is contained in the above italicized question. Since time-asymmetry cannot be explained by the time-reversal invariant laws (*equations*) of physics, it should be explained by some time-asymmetric *initial conditions*. But, at first sight, initial conditions are arbitrary; therefore, it is impossible to formulate a physical law on initial conditions. However, the initial conditions of any process are produced by previous processes in such a way that all processes in a connected universe are coordinated in some way. Therefore, the reason of time-asymmetry is the asymmetry of the universe, namely, a *global* reason. The aim of this paper is to explain this fundamental fact in a way as conceptually clear as possible.

In the early nineteenth century origin of statistical mechanics, Boltzmann asserted: "The universe, or at least a big part of it around us, considered as a mechanical system, began in a very improbable state and it is now also in a very improbable state. Then if we take a smaller system of bodies, and we isolate it instantaneously from the rest of the world, in principle this system will be in an improbable state and, during the period of isolation, it will evolve towards more probable states" [9]. Since Boltzmann's seminal work, many authors had the intuition that time-asymmetry has a global origin. For instance, Feynman claimed: "For some reason, the universe at one time had a very low entropy for its energy content, and since then entropy has increased. So that is the way towards future. That is the origin of all irreversibility" [10]. However, these traditional positions define time-asymmetry in terms of entropy increase (see

¹There are two exceptions:

a.- The second law of thermodynamics: entropy grows. But we use to consider this "law" as an empirical fact that must be demonstrated from more primitive and elementary laws.

b.- Weak interactions. But they are so weak that it is difficult to see how the time-asymmetry of the universe can be derived from these interactions. Therefore, as it is usual in the literature, we do not address this problem in this paper.

also Davies [11]). In the present paper we will reject the traditional entropic approach, following John Earman's [12] "Time Direction Heresy" according to which the arrow of time is an intrinsic, geometrical feature of space-time: this geometrical approach to the problem of the arrow of time has conceptual priority over the entropic approach since the geometrical properties of the universe are more basic than its thermodynamic properties.

In order to develop our arguments, we will adopt two methodological hypotheses which are unquestioningly accepted in present-day cosmology:

- 1.- The laws of physics are always and everywhere the same.
- 2.- The universe is unique: disconnected universes are not considered.

The paper is organized as follows:

In Section II the concepts of time-reversal invariance, irreversibility and time-asymmetry will be distinguished. Section III is devoted to introduce the difference between conventional and substantial arrows of time: this will allow us to formulate the problem in precise terms. In Section IV we will argue that the time-orientability of space-time and the existence of a global time are necessary conditions for defining the arrow of time. On this basis, we will develop a two steps program: Section V is devoted to show that the universe is a time-asymmetric object; in Section VI we will argue that this global time-asymmetry manifests itself in each point of the manifold as a local time-asymmetry. In Sections VII and VIII we will consider time-asymmetries of different origins (quantum, electromagnetic, thermodynamic, etc.), namely, other "arrows of time": since the subject of these sections is quite long but we would like to give a comprehensive overview, in many cases we will refer the reader to the literature (mostly to our own papers). In Section IX we will draw our conclusions. An Appendix studying a 1999 paper of L. Schulman [13] completes this work: the aim of this appendix is to demonstrate that the discussion on this fundamental subject is still open.

II. TIME-REVERSAL INVARIANCE, IRREVERSIBILITY, TIME-SYMMETRY

In general, these three concepts are invoked in the treatment of the problem of the arrow of time, but usually with no elucidation of their precise meanings; this results in confusions that contaminate many interesting discussions. For this reason, we will start from providing some necessary definitions:

1.- **Time-reversal invariance:** It is a property of evolution equations (laws) and, a fortiori, of the set of its solutions (models). An evolution equation is *time-reversal invariant* if it is invariant under the transformation $t \rightarrow -t$; as a result, for each solution $f(t)$, $f(-t)$ is also a solution.

2.- **Reversibility:** It is a property of a single solution of an evolution equation. A solution is *reversible* if it corresponds to a closed curve in phase space.

Both properties can combine in the four possible alternatives:

a.- **Time-reversal invariance and reversibility.** Let us consider the harmonic oscillator with Hamiltonian:

$$H = \frac{1}{2} p^2 + \frac{1}{2} K^2 q^2$$

The equation is time-reversal invariant, namely, it is invariant under the transformation $q \rightarrow q, p \rightarrow -p$. As a result, the set of trajectories is symmetric with respect to the q -axis. Since each trajectory is a closed ellipse, it is reversible.

b.- **Time-reversal invariance and irreversibility.** Let us consider the pendulum with Hamiltonian:

$$H = \frac{1}{2} p_\theta^2 + \frac{K^2}{2} \cos \theta$$

Again the equation is time-reversal invariant, and the set of solutions is symmetric with respect the θ -axis. The trajectories within the separatrices are reversible since they are closed curves. But the trajectories above (below) the separatrices are irreversible since, in the infinite time-limit, $\theta \rightarrow \infty$ ($\theta \rightarrow -\infty$). The trajectories in the upper (lower) separatrix are also irreversible since they tend to $\theta = \frac{\pi}{2}, p_\theta = 0$ ($\theta = -\frac{\pi}{2}, p_\theta = 0$) when $t \rightarrow \infty$ ($t \rightarrow -\infty$).

c.- **Time-reversal non-invariance and reversibility.** Let us consider the modified oscillator with Hamiltonian:

$$H = \frac{1}{2} p^2 + \frac{1}{2} K(p)^2 q^2$$

where $K(p) = K^+ = \text{const.}$ when $p \geq 0$, and $K(p) = K^- = \text{const.}$ when $p < 0$. If $K^+ \neq K^-$, the equation is not time-reversal invariant: the set of solutions is clearly asymmetric with respect to the q -axis. Nevertheless, each trajectory is closed and, therefore, reversible.

d.- **Time-reversal non-invariance and irreversibility.** Let us consider a damped oscillator whose equation is:

$$\ddot{q} = -K^2 q - A^2 \dot{q}$$

The equation is time-reversal non-invariant. The origin is an attractor and the trajectories are spirals: they are trapped by the origin when $t \rightarrow \infty$. Therefore, each trajectory is irreversible.

Now we will introduce what we consider the most relevant concept in the problem of the arrow of time:

3.- Time-symmetry: It is a property of a single solution of an evolution equation. A solution $f(t)$ is *time-symmetric* if there is a time t_S such that $f(t_S + t) = f(t_S - t)$.

Since we will define the arrow of time as a global feature of the universe, it is interesting to see how this definition of time-symmetry applies to cosmology. Let us consider a FRW Big Bang-Big Crunch universe ($k = 1$), where a_{\max} corresponds to a time t_{\max} ; this universe is time-symmetric if²:

$$a(t_{\max} + t) = a(t_{\max} - t)$$

More generally, let us consider a universe endowed with a cosmic time t , namely, with metric:

$$ds^2 = dt^2 - h_{ij}(t, x^k) dx^i dx^j$$

We will say that this universe is time-symmetric if there is a time t_S such that:

$$h_{ij}(t_S + t, x^i) = h_{ij}(t_S - t, x^i)$$

III. CONVENTIONAL VS. SUBSTANTIAL ARROWS OF TIME

Traditional discussions around the problem of the arrow of time are usually subsumed under the label "the problem of the direction of time", as if we could find an exclusively physical criterion for singling out *the* direction of time, identified with what we call "the future". But there is nothing in (local) physics that distinguishes, in a non-arbitrary way, between past and future as we conceive them. It might be objected that physics implicitly assumes this distinction with the use of asymmetric temporal expressions, like "future light cone", "initial conditions", "increasing time", and so on. However this is not the case, and the reason can be understood on the basis of the distinction between the concepts of conventional and substantial.

Two objects are *formally identical* when there is a permutation that interchanges the objects but does not change the properties of the system to which they belong or in whose description they are involved. In physics it is usual to work with formally identical objects: the two lobes of a light cone, the two spin senses, etc.

i.- We will say that we establish a *conventional* difference when we call two formally identical objects with two different names, e.g., when we assign different signs to the two spin senses.

ii.- We will say that the difference between two objects is *substantial* when we give different names to two objects which are not formally identical (see [14], [15]). In this case, even though the names are conventional, the difference is substantial. E.g. the difference between the two poles of the theoretical model of a magnet is conventional since both poles are formally identical; the difference between the two poles of the Earth is substantial because in the north pole there is an ocean and in the south pole there is a continent (and the difference between ocean and continent remains substantial even if we conventionally change the names of the poles).

Once this point is accepted, the problem cannot yet be posed in terms of singling out the future direction of time: the problem of the arrow of time becomes the problem of finding a *substantial difference* between the two temporal directions. But if this is our central question, we cannot project our independent intuitions about past and future for solving it without begging the question. If we want to address the problem of the arrow of time from a perspective purged of our temporal intuitions, we must avoid the conclusions derived from subtly presupposing time-asymmetric notions. As Huw Price [16] claims, it is necessary to stand at a point outside of time, and thence to regard reality in atemporal terms. This atemporal standpoint prevents us from using the asymmetric temporal expressions of our natural language in a non-conventional way. But then, what does "the arrow of time" mean when we accept this constraint? Of course, the traditional expression coined by Eddington has only a metaphorical sense: its meaning must be understood by analogy. We recognize the difference between the head and the tail of an arrow on the basis of its geometrical properties; therefore, we can substantially distinguish between both directions, head-to-tail and

²Now we are just considering the geometry. Matter within the universe will be considered in the next sections.

tail-to-head, independently of our particular perspective. Analogously, we will conceive the problem of the arrow of time in terms of *the possibility of establishing a substantial distinction between the two directions of time on the basis of exclusively physical arguments*.

When the problem is formulated in these terms, it is clear that the solution will consist in demonstrating the time-asymmetry of the universe: in a time-asymmetric universe, the two directions of time can be substantially distinguished. However, first it is necessary to explain the topological conditions required to establish such a distinction.

IV. TIME-ORIENTABILITY AND COSMIC TIME

Earman [12] and Grünbaum [17] were the first authors who emphasized the relevance of time-orientability to the problem of the arrow of time. In fact, general relativity considers the universe as a pseudo-Riemannian manifold that may be time-orientable or not. A space-time is *time-orientable* if and only if there exists a *continuous* non-vanishing time-like vector field globally defined. By means of this field, the set of all light semi-cones (lobes) of the manifold can be split into two equivalence classes, C_+ and C_- : the lobes of C_+ contain the vectors of the field and the lobes of C_- do not contain them. If space-time were not time-orientable, the distinction between future lobes and past lobes would not be univocally definable on a global level. On the other hand, in a time-orientable space-time, if there were a time-reversal non-invariant law L defined in a *continuous* way all over the manifold, this would allow us to choose one of the classes as the future class (say C_+) and the other one as the past class (say C_-); the law L would be sufficient for defining the arrow of time for the whole universe (namely, a future lobe $C_+(x)$ and a past lobe $C_-(x)$ at each point x). In fact, if one lobe of the class C_+ were considered as the future lobe at a point x and another lobe of the same class were considered as the past lobe at a point y , then joining these two points by means of a continuous curve (because we only consider connected universes) and propagating the lobe of x towards y (and vice versa) would be sufficient for finding a point where the law L would be discontinuous, contrary to our assumption.

But, which is this global continuous time-reversal invariant law that allows us to define past and future? This is the essence of Matthews' criticism [18] to the relevance of time-orientability: since there are not continuous and global time-reversal non-invariant laws of nature (but anyway the arrow of time does exist), time-asymmetry is necessarily defined by local laws and, then, it is just a local property; therefore, nothing rules out the possibility that the arrow of time points to different directions in different regions of space-time (see also Reichenbach [19]). What Matthews has forgotten is that an asymmetric physical fact can be used to define time-asymmetry instead of a physical law. Of course, it must be an ubiquitous physical fact, because it must be used to define the future and the past lobes at all the points of the universe. This ubiquitous physical fact is the global time-asymmetry of the universe as a whole, which manifests itself in all local domains (as we will explain in the next sections).

However, in order to obtain the arrow of time, the possibility of defining a cosmic time is a further requirement. A space-time has a *cosmic time* if the following conditions hold:

- i.- The space-time satisfies the *stable causality condition* [20]: in this case, the manifold M possesses a *global time function* $t : M \rightarrow \mathbb{R}$ whose gradient is everywhere time-like.
- ii.- The global time function t can be computed as the *distance* between two hypersurfaces of the resulting foliation, measured along any trajectory orthogonal to the foliation. In this case, the time function t is the cosmic time.

Under the assumptions of time-orientability and existence of a cosmic time (conditions satisfied by usual present-day cosmological models), in the following sections we will show that:

- 1.- Time-symmetric universes belong to a set of *measure zero* on the space of all possible universes³.
- 2.- The global time-asymmetry of the universe can be used *locally* at each point x to define the future and the past lobes, $C_+(x)$ and $C_-(x)$.

We will develop these points in the next two sections.

³Precisely, we refer to the Liouville measure and to any measure absolutely continuous with respect to it. In some cases it would more rigorous to say that the dimension of the subspace of time-symmetric solutions of the universe evolution is smaller than the dimension of the space of solutions, meaning that time-asymmetric solutions are "generic".

A. The theorem

In his interesting book, Price [16] emphasizes that time-reversal invariance is not an obstacle to construct a time-asymmetric model of the universe: a time-reversal invariant equation may have time-asymmetric solutions⁴. He illustrates this point with the familiar analogy of a factory which produces equal numbers of left-handed and right-handed corkscrews: the production as a whole is completely unbiased, but each individual corkscrew is asymmetric. Price argument shows the possibility of describing time-asymmetric universes by means of time-reversal invariant laws. But, what is the reason to suppose that time-asymmetric universes have high probability? We will demonstrate that time-asymmetric solutions of the universe equations have measure zero in the corresponding phase space.

Let us consider some model of the universe equations. All known examples have the following two properties (e.g. see [21], but there are many other examples):

- 1.- They are time-reversal invariant, namely, invariant under the transformation $t \rightarrow -t$.
- 2.- They are time-translation invariant, namely, invariant under the transformation $t \rightarrow t + \text{const.}$ ⁵ (homogeneous time).

Let us consider the following cases:

- a) The particular case of a FRW model where a is the only dynamical variable. It must satisfy the Hamiltonian constraint:

$$H(a, \dot{a}) = 0$$

namely, the Einstein equation. Then:

$$\ddot{a} = f(a)$$

If $f(a)$ has a root at a_S corresponding to the time t_S and $f'(a_S) \neq 0$ (which would be a generic case), then at $t = t_S$ the radius a has a maximum or a minimum. Since, by hypothesis, the equations are time-reversal invariant and homogeneous (which allows us to begin our reasoning from any point), both sides of $a(t)$ will be symmetric with respect to t_S :

$$a(t_S + t) = a(t_S - t)$$

Therefore, for this kind of universe global time-symmetry is generic⁶ to the extent that the dimension of the set of time-symmetric solutions is the same than the dimension of phase space.

- b) Let us now consider a more generic case: a FRW universe with radius a and matter represented by a neutral scalar field ϕ . The dynamical variables are now $a, \dot{a}, \phi, \dot{\phi}$. They satisfy a generic Hamiltonian constraint:

$$H(a, \dot{a}, \phi, \dot{\phi}) = 0 \tag{1}$$

which reduces the dimension of phase space from 4 to 3; then, we can consider a phase space of variables $\dot{a}, \phi, \dot{\phi}$ and:

$$a = f(\dot{a}, \phi, \dot{\phi}) \tag{2}$$

a function obtained solving eq.(1).

⁴Of course, this "loophole" is not helpful when we are dealing with a multiplicity of systems: for each time-asymmetric solution there is another time-asymmetric solution that is the temporal mirror image of the first one. But when we are studying the whole universe, both solutions are equivalent descriptions of one and the same universe.

⁵We are referring to the equations that rule the behavior of the universe, not to the particular solutions that normally do not have time-translation symmetry.

⁶Essentially this was what Hawking called his "greatest mistake" [22]. In fact, time-symmetry appears to be generic only in the simplest cosmological universes.

If we want to obtain a time-symmetric continuous⁷ solution such that $a \geq 0$ ⁸, there must be a time t_S regarding to which a is symmetric:

$$a(t_S + t) = a(t_S - t) \quad \text{and} \quad \dot{a}(t_S) = 0$$

In order to obtain complete time-symmetry, ϕ must also be symmetric about t_S . There are two cases: even symmetry:

$$\phi(t_S + t) = \phi(t_S - t) \quad \text{and} \quad \dot{\phi}(t_S) = 0$$

and odd symmetry:

$$\phi(t_S + t) = -\phi(t_S - t) \quad \text{and} \quad \phi(t_S) = 0$$

This means that time-symmetric trajectories necessarily pass through the axes $(0, \phi, 0)$ or $(0, 0, \dot{\phi})$ of the phase space. From these "initial" conditions we can propagate, using the evolution equations, the corresponding trajectories; this operation will produce two surfaces that contain the trajectories with at least one point of symmetry, that is, that contain all the possible time-symmetric trajectories. Both surfaces have dimension $2 < 3$ (namely, the dimension of our phase space). The usual Liouville measure of these sets is zero, and also any measure absolutely continuous with respect to it. In this way we have proved that, for generic models of the universe, the solutions are time-asymmetric with the exception of a subset of solutions of measure zero. q.e.d.⁹.

B. Generalization of the theorem

The theorem can be easily generalized to the case where ϕ has many components, or to the case of many fields with many components. Some of these fields may be fluctuations of the metric: in this case, we must Fourier transform the equations, and this would allow us to reproduce the theorem only with t functions. Since properties 1 and 2 (time-reversal invariance and time-translation invariance) are also true in the classical statistical case, the theorem can be also demonstrated in this case¹⁰. And also in the quantum case, albeit some quantum gravity problems like time definition [23].

Let us now consider the coarse-grained version of the theorem. Let ε be the size of the grain and, in order to compare measures, let us consider that the phase space $\dot{a}, \phi, \dot{\phi}$ is a cube of volume L^3 . In this case, boundary conditions $(0, \phi, 0)$ and $(0, 0, \dot{\phi})$ will be fuzzy and the volume of the set of time-symmetric initial conditions will have measure $2\varepsilon^2 L$. This magnitude can be compared with the size of the phase space, obtaining the ratio $2\varepsilon^2 L / L^3 = 2(\varepsilon/L)^2$. Of course, in the usual case $\varepsilon \ll L$; then, the measure of the set of points corresponding to initial conditions that lead to time-symmetric universes is extremely smaller than the measure of the phase space. The same argument can be applied to the set of time-symmetric solutions with measure $2\varepsilon L^2$, where $2\varepsilon L^2 / L^3 = 2\varepsilon/L \ll 1$ if $\varepsilon \ll L$. q.e.d.

This completes the first argument announced in Section IV; let us now develop the second argument.

VI. FROM GLOBAL TIME-ASYMMETRY TO LOCAL TIME-ASYMMETRY

A. The generic case

Combining the results of Section IV and Section V it is not difficult to achieve the second step of our program.

⁷We will disregard non-continuous solutions since normally information do not pass through discontinuities and we are only considering *connected* universes where information can go from a point to any other timelike connected point.

⁸As only a^2 appears in a FRW metric, we will consider just the case $a \geq 0$ since the point $a = 0$ is actually a singularity that cuts the time evolution.

⁹Of course, the models must be generic. E.g. if we chose an H such that $\frac{\partial H}{\partial \phi} = \frac{\partial H}{\partial \dot{\phi}} = 0$, we are in the case of point a) and time-symmetric solutions become generic. But this is not a generic case. A more (apparently) non-generic example could be obtained making the canonical transformation of the model above.

¹⁰When the phase space has infinite dimensions, it is better to use the notion of dimension instead of that of measure, as explained in footnote 3.

From Section IV we know that if a space-time is time-orientable, a continuous non-vanishing time-like vector field $\gamma^\mu(x)$ can be defined all over the manifold. At this stage, the universe is *time-orientable* but not yet *time-oriented*, because the distinction between $\gamma^\mu(x)$ and $-\gamma^\mu(x)$ is just conventional. Now Section V comes into play. A time-orientable space-time having a cosmic time t is time-asymmetric if there is not a time t_S that splits the manifold into two "halves", one the temporal mirror image of the other regarding their intrinsic geometrical properties. This means that, in a time-asymmetric universe, any time t_A splits the manifold into two sections that are different to each other: the section $t > t_A$ is *substantially* different than the section $t < t_A$. We can choose any point x_0 with $t = t_A$ and conventionally consider that $-\gamma^\mu(x_0)$ points towards $t < t_A$ and $\gamma^\mu(x_0)$ points towards $t > t_A$ or vice versa: in any case we have established a substantial difference between $\gamma^\mu(x_0)$ and $-\gamma^\mu(x_0)$. We can conventionally call "future" the direction of $\gamma^\mu(x_0)$ and "past" the direction of $-\gamma^\mu(x_0)$ or vice versa, but in any case past is substantially different than future. Now we can extend this difference to the whole continuous fields $\gamma^\mu(x)$ and $-\gamma^\mu(x)$: in this way, the time-orientation of the space-time has been established. Since the field $\gamma^\mu(x)$ is defined all over the manifold, it can be used *locally* at each point x to define the future and the past lobes: for instance, if we have called "future" the direction of $\gamma^\mu(x)$, $C_+(x)$ contains $\gamma^\mu(x)$ and $C_-(x)$ contains $-\gamma^\mu(x)$. This is the solution of the second step in the general case.

B. From the generic case towards our own universe.

Even if the second step has been completed in the generic case, it is clear that the solution is more mathematical than physical. Thus, it would be desirable to show how the general time-orientation is reflected in everyday physics, where time-asymmetry manifests itself in terms of time-asymmetric energy fluxes. However, this task will lead us to impose reasonable restrictions in the considered cosmological model in such a way that the explanation of local time-asymmetry applies, not to the generic case, but rather to the particular case of our own universe.

i.- Up to this point, global time-asymmetry has been considered as a substantial asymmetry of the geometry of the universe, embodied in the metric tensor defined at each point of the space-time: $g_{\mu\nu}(x)$. Perhaps the easiest way to see how this geometrical time-asymmetry is translated into local physical terms is to consider the energy-momentum tensor, which can be computed by using $g_{\mu\nu}(x)$ and its derivatives through Einstein's equation:

$$T_{\mu\nu} = \frac{1}{8\pi} \left(R_{\mu\nu}(g) - \frac{1}{2} g_{\mu\nu} R(g) - \Lambda g_{\mu\nu} \right)$$

The curvatures $R_{\mu\nu}(g)$, $R(g)$ can be obtained from $g_{\mu\nu}(x)$ and its derivatives, and Λ is the cosmological constant. Now we impose a first condition: that our $T_{\mu\nu}$ turns out to be a "normal" or *Type I* energy-momentum tensor. Then, $T_{\mu\nu}$ can be written as:

$$T_{\mu\nu} = s_0 V_\mu^{(0)} V_\nu^{(0)} + \sum_{i=1}^3 s_i V_\mu^{(i)} V_\nu^{(i)}$$

where $\{V_\mu^{(0)}, V_\mu^{(i)}\}$ is an orthonormal tetrad, $V_\mu^{(0)}$ is time-like and the $V_\mu^{(i)}$ are space-like ($i = 1, 2, 3$) (see [20], [24]). Since we have assumed that the manifold is continuous, $g_{\mu\nu}(x)$ and $T_{\mu\nu}(x)$ are continuously defined over the manifold (provided the derivatives of $g_{\mu\nu}(x)$ are also continuous); this means that $V_\mu^{(0)}(x)$ is a continuous unitary time-like vector field defined all over the manifold, which can play the role of the field $\gamma^\mu(x)$ if everywhere $s_0 \neq 0$ ($V_\mu^{(0)}$, even if time-like, may change its sign when $s_0 = 0$).

Here we impose a second condition: that the universe satisfies the *dominant energy condition*: i.e. $T^{00} \geq |T^{\mu\nu}|$ in any orthonormal basis (namely, $s_0 \geq 0$ and $s_i \in [-s_0, s_0]$). In this case, $s_0 \neq 0$ and, then, $V_\mu^{(0)}(x)$ is continuous, time-like and non-vanishing. This means that $V_\mu^{(0)}(x)$ can play the role of $\gamma^\mu(x)$, with the advantage that it has a relevant physical sense. In this way, a time-orientation is chosen at each point x of the manifold, and the time components of $T_{\mu\nu}$ acquire definite signs according to this orientation. We will use this orientation in our arguments below. Therefore we have translated the global time-asymmetry into local terms, endowing the local arrow with a physical sense.

ii.- Since we are now in local grounds, our new task is to understand the *local nature* of the characters in the play. If $T^{00} \geq |T^{\mu\nu}|$, then $T^{00} \geq |T^{i0}|$. Therefore, $T^{0\mu}$, which is usually considered as the local energy flux, is a time-like (or light-like) vector. This holds for all presently known forms of energy-matter and, so, there are in fact good reasons

for believing that this should be the case in almost all situations (for the exceptions, see [25]¹¹).

iii.-But, is really $T^{0\mu}$ the energy flux? To go even closer to everyday physics, we must remember that $T_{\mu\nu}$ satisfy the "conservation" equation:

$$\nabla_\mu T^{\mu\nu} = 0$$

Nevertheless, as it is well known, this is not a true conservation equation since ∇_μ is a covariant derivative. The usual conservation equation with ordinary derivative reads:

$$\partial_\mu \tau^{\mu\nu} = 0$$

where $\tau_{\mu\nu}$ is not a tensor and it is defined as:

$$\tau_{\mu\nu} = \sqrt{-g} (T_{\mu\nu} + t_{\mu\nu})$$

where we have introduced a $t_{\mu\nu}$ that reads:

$$\sqrt{-g} t_{\mu\nu} = \frac{1}{16\pi} \left[\mathcal{L} g_{\mu\nu} - \frac{\partial \mathcal{L}}{\partial g_{\mu\nu, \lambda}} g_{\mu\nu, \lambda} \right]$$

where \mathcal{L} is the Lagrangian. $t_{\mu\nu}$ is also an homogeneous and quadratic function of the connection $\Gamma_{\nu\mu}^\lambda$ [27]. Now we can consider the coordinates $\tau^{0\mu}$, which satisfy:

$$\partial_\mu \tau^{0\mu} = \partial_0 \tau^{00} + \partial_i \tau^{0i} = 0$$

namely, a usual conservation equation. Even if $\tau^{0\mu}$ is not a four-vector, it can be defined in each coordinate system: in each system τ^{00} can be considered as the energy density and τ^{0i} as the energy flux (the Poynting vector). This means that the field $\tau^{0\mu}(x)$ represents the spatio-temporal energy flow within the universe better than $T^{0\mu}$.

In particular, in a local inertial frame where $\Gamma_{\nu\mu}^\lambda = 0$, we have $\tau_{\mu\nu} = \sqrt{-g} T_{\mu\nu}$: in orthonormal coordinates, the dominant energy condition will be now $\tau^{00} \geq |\tau^{i0}|$ and $\tau^{0\mu}$ will be time-like (or light-like). But $\tau^{0\mu}$ is just a local energy flow since it is defined in orthonormal local inertial frames. Nevertheless, in any moving frame with respect to those ones, if the acceleration of the moving frame is not very large, the $(\Gamma_{\nu\mu}^\lambda)^2$ and the $t_{\mu\nu}$ are very small and the energy flux in the moving frame is time-like (or light-like) for all practical purposes. This is precisely the case of the commoving frame of our present-day universe: the arrows of Fig.1 represent the $\tau^{0\mu}$ in this case.

In summary, $\tau^{0\mu}$ (that can locally be considered as the four-velocity of a quantum of energy carrying a message) is a time-like *local* energy flux and:

- a) It inherits the global time-asymmetry of $g_{\mu\nu}(x)$, i.e., the geometrical time-asymmetry of the universe.
- b) It translates the global time-asymmetry into the local level: the lobes $C_-(x)$ receive an incoming flux of energy while the lobes $C_+(x)$ emit an outgoing flux of energy.

iv.- In order to complete the argument, we should consider all possible universes, that is, not only our (Big Bang-Big Chill) universe, but also all conjectural universes (satisfying the conditions required above). Of course, in practice this is an impossible task to the extent that we do not know the phenomenology of these conjectural universes. Therefore, in the next subsection we will continue the analysis of the energy flux just in the universe we inhabit.

C. Particular case: our own universe. The Reichenbach-Davies diagram

In his classical book about the arrow of time, Hans Reichenbach [19] defines the future direction of time as the direction of the entropy increase of the majority of branch systems, that is, systems which become isolated or quasi-isolated from the main system during certain period. Paul Davies [11] appeals to Reichenbach's notion, claiming that branch systems emerge as the result of a chain or hierarchy of branchings which expand out into wider and wider regions of the universe; therefore "the origin of the arrow of time refers back to the cosmological initial conditions".

¹¹E.g., some exceptions are: Casimir effect, squeezed vacuum, Hawking evaporation, Hartle-Hawking vacuum, negative cosmological constant, etc. These objects are strange enough in nowadays observational universe to exclude the practical existence of zones with T^{00} of different signs and, therefore, with different time directions (see also [26]).

On the basis of this idea, in previous papers we have introduced the "Reichenbach-Davies diagram" [5], [7], [28]¹², where all the local processes which go from non-equilibrium to equilibrium are connected in such a way that the "output" of a process is the "input" of another one: the energy provided by a process relaxing to equilibrium serves to drive another process to non-equilibrium. This "cascade" of processes defines a global energy flux in our universe which, if traced back, owes its origin to the initial global instability that is the source of all the energy of the universe. The existence of the initial global instability can be deduced from the equations of Section V. If we consider that the universe begins in a Big Bang with $a = 0$, from eq.(2) all possible Big Bang initial conditions are contained in the surface:

$$f(\dot{a}, \dot{\phi}, \dot{\phi}) = 0$$

of phase space. The generic points of this surface are unstable (i.e. they are not attractors); then, a generic Big Bang beginning is unstable. Of course, the unstable nature of the solutions of the generic Einstein equation is a well known fact. Moreover, using entropic considerations the initial instability is studied in [29] and bibliography therein. Quantum universes are also unstable, e.g. Vilenkin universe, which is essentially a quantum system with a potential barrier allowing quantum tunneling [30]. Another unstable quantum model is studied in [31], etc. This means that the existence of the initial instability is well established.

The Reichenbach-Davies global system is the system of all time-asymmetric processes within the universe: any process of the system begins in an unstable state that was produced using energy coming from another process of the global system¹³. In fact, energy always comes from unstable→stable or non-equilibrium→equilibrium processes (coal burning into ashes, H burning into He, etc.). The global system (symbolized in Fig.1) has a time-asymmetry that we may call "*global arrow of time*" (GAT): this arrow points in opposite direction to the initial cosmological instability and follows the evolution of all the hierarchical chain towards equilibrium¹⁴. Each system in the diagram is called a *branch system* and it is represented by a box. The arrows coming from the left of each box represent the energy produced by other boxes: part of the energy going to the right is used to produce new unstable states, and the rest is degraded (the degraded energy is represented by the outgoing arrows that do not end in a box). This is also an asymmetry of the diagram of Fig.1: the arrows corresponding to degraded energy only appear at the right of each box. This is much more than just a detail, since it means that we have a concentrated source of energy in the extremity of the universe that we call "the past", from which we can pump energy to create other concentrated sources of energy, while energy gets diluted towards what we call "the future". In fact, this is the case of realistic Big Bang-Big Chill cosmological models: energy is concentrated towards the past and diluted towards the future [33], and this is another manifestation of the time-asymmetry of the universe.

The global system also allows us to introduce the notion of *causality* in the universe (i.e. *global causality*) since, using Fig.1, we can say that events A and B are not causally related, while C is the (partial) cause of D, and D is the (partial) effect of C. On this basis we can say that *no effect can occur before its cause*: this statement is meaningful because we have a global time-asymmetry that defines the word "before". The physical substratum of causality is the energy coming from an unstable state and creating a new unstable state; unstable structures are created by pumping energy from sources in the past and decay spontaneously to equilibrium towards the future.

Now we can go back to the manifold description of the previous subsection. As we have seen, the time-asymmetry of the global energy flux is a consequence of the time-asymmetry of the geometry of the universe. The direction of the energy flux on a time-orientable space-time defines a global *time-orientation*: the incoming flux defines the lobe $C_-(x) \in C_-$ at each point x , the outgoing flux defines the lobe $C_+(x) \in C_+$ at x , and all the lobes of class C_- point towards the initial instability. In this way, the global time-asymmetry of the universe defines the local time-asymmetry

¹²At the classical level, the "Reichenbach-Davies diagram" can be considered as the combination of all the classical scattering processes within the universe. We have called "Reichenbach-Bohm diagram" to the combination of all the quantum scattering processes within the universe, i.e. the quantum version of the "Reichenbach-Davies diagram" [4], [6].

¹³Since Reichenbach [19] does not take into account the time-orientability of space-time, he accepts the possibility of a universe with no global arrow of time, with regions having arrows "pointing" to opposite directions. However, this possibility not only can be objected on theoretical basis (see Section IV), but also seems unpalatable on observational grounds: nowadays we know that visible universe has a unique arrow of time, since there is no astronomical observation showing that the time-asymmetric behavior of nature would be inverted in some (eventually very distant) regions of the universe. In fact, supernovae evolutions always follow the same pattern (from birth to death, like human beings !!!), and there is no trace of an inverted pattern in all the universe. This is a relevant observation when we consider that supernovae are the markers used to measure the longest distances corresponding to objects near the limit of the visible universe [32].

¹⁴At least in an expanding universe (Big Bang-Big Chill) case, which seems to be the case of our universe.

in each one of its points. This means that the energy flux is *the ubiquitous phenomenon* that locally defines the arrow of time, because it can be found everywhere in the universe. If two sections of the universe are not connected by such a flux, then they are completely isolated from each other, and each one of them can be conceived as a universe by itself; but this situation is not considered, according to the second methodological hypothesis of the Introduction. The global Reichenbach-Davies diagram defines the arrow of time of our universe and, in this scenario, we can see how unstable states reach equilibrium becoming stable states, and how entropy grows since it grows from unstable to stable states. Therefore, the different "arrows of time" (cosmological, thermodynamic, quantum, electromagnetic, etc.) can be coordinated, as we will see in the next sections (see also [5]).

VII. OTHER CLASSICAL ARROWS OF TIME

The time-asymmetry studied in the previous sections reappears in the phenomena explained by many chapters of physics, giving rise to many "arrows of time" corresponding to the different chapters, e.g.: the cosmological arrow of time (CAT), the electromagnetic arrow of time (EMAT), the quantum arrow of time (QAT), the thermodynamic arrow of time (TAT), etc. They all result from the global time-asymmetry (GAT) and, therefore, point to the same direction. In this section we will only give a schematic introduction to this subject, since its full treatment exceeds the limits of this paper; however, it is necessary to mention these points in order to supply a complete account of the problem of the arrow of time. For each arrow (with the exception of CAT) we will show that the time-reversal invariant equations of evolution always produce two mathematical structures symmetrically related by a time-reversal transformation (that we will call "*t-symmetric twins*"), which usually embody notions related with irreversibility. However, at this level the two structures are only conventionally different: the problem here consists in supplying a non-arbitrary criterion for choosing one of them. Only GAT allows us to select one of these structures (one of the twins) as the one related to the future or as the physically relevant one for spontaneous evolutions, by creating a substantial difference between them.

A. Cosmological arrow of time

CAT is embodied in the fact that the radius of the universe grows, using the direction of time defined by GAT (if not, the sentence would be meaningless). Then, CAT points to the same direction than GAT in expanding universes like ours. The usual irreversible models of present-day cosmology ($a \sim t^{\frac{1}{2}}$ or $a \sim t^{\frac{2}{3}}$ or $a \sim e^{Ht}$) are all growing, and recent cosmological observations show that this is the case for our actual universe [34]. Let us observe that even if CAT changes its direction when a (conjectural) universe passes from an expanding phase to a contracting phase, GAT never changes since it is defined by the global time-orientation of the universe-manifold, as explained in Section VI.

B. Electromagnetic arrow of time

i.- EMAT is embodied in the fact that we must use retarded solutions in electromagnetic problems. But electromagnetism provides us a pair of advanced and retarded solutions (the pair of *t-symmetric twins* corresponding to this case) which are only conventionally different. These solutions are related also with the incoming and outgoing states in scattering situations. The mathematical structure of these incoming and outgoing states is rigorously defined by the Lax-Phillips scattering theory [35].

ii.- In the Reichenbach-Davies diagram of Fig.1 we see that all spontaneous evolutions (namely, the arrows going towards the right side of the diagram) are outgoing states from the corresponding boxes, and these outgoing (spontaneously evolving) states correspond to retarded solutions¹⁵. The incoming states (the arrows going into the boxes coming from the left side) are not spontaneous evolutions since they pump energy from the past to create unstable states, and correspond to advanced solutions. Therefore, the Reichenbach-Davies diagram, which defines the past and the future directions of time, generates also a substantial difference between both members of the pair of *t-symmetric*

¹⁵EMAT can also be explained on cosmological grounds by the existence of an absorbing black body at decoupling time [36]. Moreover, in Lax-Phillips theory outgoing states correspond to retarded solutions and to Hardy classes from below, anticipating the results of Subsection VIII.B.

twins: retarded solutions are obtained by means of energy coming from the past, whereas advanced solutions are obtained by means of energy coming from the future. Since the global energy flux comes from the past initial instability, only retarded solutions are physically meaningful: this fact defines EMAT which, so defined, coincides with GAT.

C. Thermodynamic arrow of time

i.-TAT corresponds to the entropy increase in spontaneous evolutions of closed systems, as prescribed by the second law of thermodynamics. It is well known that many processes within the universe are chaotic: the existence of chaotic mixing systems (like the well known examples of the Gibbs ink drop, the sugar lump or the bottle of perfume) is an obvious fact. Chaotic subsystems within the universe make the universe as a whole a chaotic system since some of its parts are chaotic. For spontaneous evolutions of chaotic closed systems an entropy can be defined [37], *which monotonically increases from non-equilibrium in the present to equilibrium in the future*. But, if starting from non-equilibrium in the present we compute the entropy evolution towards the past, we will see that entropy also grows in this time direction. This fact was already pointed out by Paul and Tatiana Ehrenfest [38] in their criticisms to Gibbs' approach (see also [39]). The two temporally opposed evolutions are the pair of t-symmetric twins corresponding to this case. The difference between the twins is just conventional, and both express the irreversible nature of thermodynamic processes going towards equilibrium either in the past or in the future.

ii.- To break this symmetry we must know which one of both processes corresponds to the spontaneous evolution, as we explained at the end of the previous section. Using this criterion we can know which is the spontaneous evolution and which is the non-spontaneous one. Then we can meaningfully say that entropy grows in the spontaneous evolutions of closed systems. So GAT gives rise again to a substantial difference between the two members of the pair. If we define the "total entropy" of the matter-energy within the universe¹⁶ by adding the entropies of the *spontaneous evolutions* of all the closed subsystems of the universe, since all these entropies increase *from non-equilibrium to equilibrium*, the total entropy of the universe increases in the same direction, from the initial unstable state to the final equilibrium state. Therefore, TAT points to the same direction as GAT¹⁷. Of course, we may use as well other ways of defining the "total entropy" of the universe, e.g.: we can define directly a conditional entropy and then find that this entropy increases at least during the period where the universe is close to thermal equilibrium [29], or we can consider the entropy produced by the creation of particles in the early universe [7] and obtain that it also increases.

D. Chaotic processes (from classical probabilities to classical facts)

Before we roll a (classical) dice, we can only foresee that the probability for each face to come up is 1/6 (this is a classical probabilistic prophesy, quantum ones will be treated in Subsection VIII.C) but, of course, we cannot say which is the face that will come up. After the dice is rolled, we certainly know which face is up: this a classical fact and such facts produce a classical history. We must explain why this evolution *probabilities* \rightarrow *facts* (or if you prefer *prophesy* \rightarrow *history*) occurs towards the future and not the toward the past.

i.-The dice is laying on the table at time $-T$ and we know exactly which one of its faces is up. Then, it undergoes a chaotic motion (in a dice cup and a gambling table). Finally it remains at rest on the table at time T , and we know again which one of its faces is up. This means that, under this description (even though not in all its details) the phenomenon is time-symmetric: we are certain at $t = \pm T$ and we can state just probabilities at, let us say, $t = 0$. This phenomenon can be decomposed in a pair of t-symmetric twins: the process from $t = -T$ to $t = 0$ and the process from $t = 0$ to $t = T$. Again, the problem consists in supplying a non-conventional difference between both processes,

¹⁶We are not considering an eventual entropy of the gravitational field.

¹⁷Here a simple example of a Lyapunov variable could be in order. Let us consider the typical and polemical case where a has a maximum at some t and begins and ends as $a = 0$. As in eq. (2) we can have

$$\dot{a} = -L(a, \phi, \dot{\phi})$$

and $\ddot{a} \leq 0$ for all the evolution. So L is a Lyapunov evergrowing variable that depends on matter (through $\phi, \dot{\phi}$) and on geometry (through a), as we would expect. This trivial fact shows how easy it is to find Lyapunov variables as soon as the universe has just small complexity. Another example is the "entropy gap", as studied in paper [29] and bibliography therein.

that is, in explaining why we only consider the process $t = 0 \rightarrow t = T$ and never the process $t = -T \rightarrow t = 0$ when we gamble¹⁸.

ii.- The explanation begins by considering the states at $t = \pm T$ as equilibrium states and the state at $t = 0$ as a state out of equilibrium. In fact, the dice was on the table with some face up at $t = -T$. We take the dice out of this equilibrium position at $t = 0$ by using energy (from a box of the Reichenbach-Davies diagram), and put it in chaotic motion (if we do not use energy, the dice will stay forever in its equilibrium position). Then, at $t = T$ the dice ends in its final equilibrium state. The first process is forced, since we have pumped energy from the past; the second process is spontaneous, since the dice returns spontaneously to a equilibrium position. So the global asymmetry of the Reichenbach-Diagram (i.e. GAT) introduces the desired substantial difference between the forced evolution $t = -T \rightarrow t = 0$ that belongs to $C_-(x)$ and the spontaneous evolution $t = 0 \rightarrow t = T$ that belongs to $C_+(x)$.

Now we understand why classical probabilities (prophesy) always comes before than classical facts (history). We can state this subject in another way: the Reichenbach-Davies diagram allows us to explain the history of an observer along his/her time-like path as the increase of his/her amount of information [2]. The main fact is that messages are transported by means of energy originated in unstable states that evolve towards equilibrium. Therefore, the flow of messages follows the universal flow of energy described in Section VI, namely, from past to future. This means that the information increase of an observer can be illustrated as in Fig.2, where the central curve symbolizes the observer space-time path and the curves arriving to this path are the messages carried by energy coming from unstable states (symbolized by boxes). We can then see that the amount of information of the observer increases from the left side to the right side of the diagram, creating his/her history.

VIII. THE QUANTUM ARROW OF TIME

A. Quantum measurement arrow of time.

Let us consider the simple example of a measurement process in the scattering experiment of Fig.1 (dotted box). If we want to describe the complete preparation-measurement process, we must consider not only the scattering process itself, but also the accelerator that prepares the beam, and the measurement apparatus, namely, the detector. The accelerator obtains its energy from a source, where a decaying process takes place. In the detector, a creation of an unstable state and a decaying process occur, e.g., the particles to be detected are counted by a Geiger counter, where they interact with a gas whose states are first excited (creation process) and then decay (the energy obtained from this decaying process is used to count the passing particles). The complete process of preparation-measurement corresponds to the dotted box of Fig.1, that we reproduce in Fig.3.

i.- Nevertheless, if we do not consider the origin or the fate of the energies related to the preparation-measurement process (i.e. the direction of the flux of energy), we have no substantial criterion for distinguishing between the original scattering process and its temporal mirror image: this is another example of t-symmetric twins.

ii.- But every preparation-measurement process takes place within the Reichenbach-Davies system, since the energy comes from a source of energy that can only be found in this global system. Therefore, the process that goes from preparation to measurement turns out to be essentially different from the time-reversed one: since preparation needs the energy that comes from the hierarchical chain, it is the first process; the measurement is a decaying process that produces degraded energy flowing towards the future. Since QAT goes from *preparation* to *measurement* [41], its direction agrees with the direction of GAT.

B. Minimal irreversible quantum mechanics

The time-asymmetry of the universe also leads to the possibility of formulate an irreversible quantum mechanics [42], [43] or, more precisely, a *minimal* irreversible quantum mechanics [2], [3], [6], [44].

¹⁸All aleatory phenomena in nature can be analyzed as we have done with the dice. E.g. the water in the sea is in an equilibrium state. The energy coming from the sun (that we can consider as a source of energy where H is transformed in He and, therefore, can be placed in the Reichenbach-Davies diagram) evaporates this water. Now the same water belongs to a meteorological chaotic system whose state can only be approximately foreseen (with only 4 or 5 days of anticipation). Then, rain falls and the same water ends in the sea, in an equilibrium state as initially.

i.- In fact, when we make the analytical extension of the energy spectrum of the quantum system's Hamiltonian into the complex plane, we find poles in the lower half-plane corresponding to decaying unstable states, and also symmetric poles in the upper half-plane corresponding to growing unstable states. These poles obviously correspond to irreversible processes. Moreover, the symmetric position of these poles shows the time-reversal invariance of the evolution equation: these pairs of poles can be considered as the best illustration of pairs of t-symmetric twins.

ii.- But the growing states are created by the energy pumped by previous unstable states, and the decaying states provide energy for latter processes: the first energy flux is contained in $C_-(x)$ and the second in $C_+(x)$. Therefore, the time-asymmetry of the universe defines the substantial difference between growing and decaying unstable states, and allows us to distinguish which poles are growing and which poles are decaying. A cosmological description of this fact is given in [31].

Moreover, in order to transform the poles in Gamov vectors it is necessary to introduce two subspaces of the Hilbert space \mathcal{H} , which we will call ϕ_\pm [3]. The states $|\varphi\rangle$ of the subspace ϕ_+ (ϕ_-) are characterized by the fact that their projections $\langle\omega|\varphi\rangle$ (on the energy eigenstates $|\omega\rangle$) are functions of the Hardy class from above (below). We can prove that incoming quantum states belong to the Hardy class from above and outgoing quantum states belong to the Hardy class from below, being this fact the basis to understand decaying and growing processes and, therefore, also the basis of irreversible quantum mechanics (see [6]). But, up to this point, the Hilbert space \mathcal{H} is time-reversal invariant in the sense that:

$$K\mathcal{H} = \mathcal{H}$$

where K is the Wigner antilinear time-inversion operator. However, spaces ϕ_\pm are not time-reversal invariant since:

$$K\phi_\pm = \phi_\mp$$

Then, Hilbert space is a space that cannot be used to formulate a time-reversal non-invariant quantum mechanics [5]; the substitution $\mathcal{H} \rightarrow \phi_-$ or $\mathcal{H} \rightarrow \phi_+$ is the minimal modification that we should introduce into quantum mechanics in order to make the theory time-reversal non-invariant. Nevertheless, up to this point the difference between ϕ_+ and ϕ_- is just conventional (they are t-symmetric twins). But we can establish a substantial difference if we use GAT to relate the incoming states with the flux of energy coming from the initial global instability and the outgoing states with the flux of energy that goes to the final equilibrium state of the universe. This amounts to consider the space ϕ_- as the physically meaningful and, therefore, to retain the lower poles corresponding to the spontaneous decaying processes belonging to $C_+(x)$.

C. Decoherence and the classical limit

After this two introductory subsections, we would like to study the so-called classical limit (i.e. the essential ingredient of the measurement process) in detail. This limit can be conceptually analyzed in two steps. The first step turns quantum mechanics into classical statistical mechanics; in this phase we deal with probabilities. In the second step we go from classical statistical mechanics to classical mechanics. The classical limit is one of the phenomena that more eloquently shows time-asymmetry since this limit always occurs towards the future.

Let us study the two steps in detail.

1. Quantum mechanics \rightarrow classical statistical mechanics

By means of the formalism of paper [45], we will study a quantum system with energy spectrum $0 \leq \omega < \infty$ in the simplest case where the CSCO is just H^{19} . The observable O , belonging to some space \mathcal{O} , can be expressed in terms of the Hamiltonian eigenbasis $\{|\omega\rangle\}$ as:

$$O = \int O(\omega) |\omega\rangle d\omega + \int \int O(\omega, \omega') |\omega; \omega'\rangle d\omega d\omega' \quad (3)$$

¹⁹More complete CSCO's are studied in papers [45] and [46].

where $|\omega\rangle = |\omega\rangle\langle\omega|$, $|\omega; \omega'\rangle = |\omega\rangle\langle\omega'|$, and $O(\omega, \omega')$ are regular functions such that the Riemann-Lebesgue theorem can be applied in eq.(5) [44]. Let us note that the first term of the r.h.s. of the last equation is a diagonal operator, while the second one is not diagonal. If ρ is a quantum state belonging to a convex $\mathcal{S} \subset \mathcal{O}'$, it can be expanded as:

$$\rho = \int \rho(\omega) |\omega\rangle\langle\omega| d\omega + \int \int \rho(\omega, \omega') |\omega; \omega'\rangle\langle\omega; \omega'| d\omega d\omega' \quad (4)$$

where $\{|\omega\rangle, |\omega; \omega'\rangle\}$ is the dual basis of $\{|\omega\rangle, |\omega; \omega'\rangle\}$. Again, $\rho(\omega, \omega')$ are regular functions such that the Riemann-Lebesgue theorem can be applied in eq.(5): the first term of the r.h.s. of the last equation is a diagonal operator but not the second one. Then, the time evolution of the mean value of O in the state ρ , $\langle O \rangle_\rho$, reads:

$$\langle O \rangle_\rho = \langle \rho | O \rangle = \int \rho(\omega) O(\omega) d\omega + \int \int \rho(\omega, \omega') O(\omega, \omega') e^{-i(\omega - \omega')t} d\omega d\omega' \quad (5)$$

Using Riemann-Lebesgue theorem, we have:

$$\lim_{t \rightarrow \pm\infty} \langle O \rangle_{\rho(t)} = \lim_{t \rightarrow \pm\infty} \langle \rho | O \rangle = \langle O \rangle_{\rho_*} = \langle \rho_* | O \rangle \quad (6)$$

where O is any observable of the space \mathcal{O} and ρ_* is a diagonal operator:

$$(\rho_*| = \int \rho_\omega |\omega\rangle\langle\omega| d\omega \quad (7)$$

Then, we have the *weak limit*:

$$W \lim_{t \rightarrow \pm\infty} |\rho\rangle = |\rho_*| \quad (8)$$

and we obtain decoherence. This method is used in systems with more complex energy spectra in papers [45] and [46].

i.- If the state ρ at $t = 0$ is a generic non-diagonal state, it *weakly* tends to the diagonal state ρ_* towards the past and towards the future, showing the time-reversal invariance of the underlying evolution equations. This fact means that, if the normal equilibrium state of a local system is essentially classical (namely, represented by a diagonal operator), then it begins classical when $t \rightarrow -\infty$, is taken out from this condition and becomes quantum when $t = 0$ (that is, represented by a non-diagonal density matrix), and finally tends again to a classical state when $t \rightarrow +\infty$. Analogously to the case of Subsection VII.D, this phenomenon can be decomposed in a pair of t-symmetric twins: the process from $t = -\infty$ to $t = 0$ and the process from $t = 0$ to $t = \infty$. Again, the problem consists in supplying a non-conventional difference between both processes, that is, in explaining why classicality emerges in the process $t = 0 \rightarrow t = \infty$ but we never find a classical system becoming quantum in the process $t = -\infty \rightarrow t = 0$.

ii.-However, a system undergoing decoherence is just a particular case of an unstable structure, which is created by pumping energy from the past and which decays into equilibrium towards the future²⁰. As we have seen, the Big Bang corresponds to high energies and, therefore, it is most likely a quantum state that decays into a classical universe radiating energy. This energy produces local quantum states which, in turn, decay into classical states. In this way, the universe, that begins in a quantum state, gradually becomes classical in each one of its subsystem. This is the global description of the transition from quantum mechanics to classical statistical mechanics in the universe.

Now we can go back to Subsection VIII.B in order to see how the notions of that subsection are related with decoherence. The spontaneous evolution from quantum to classical happens in $C_+(x)$ (as all the spontaneous evolutions) and, therefore, it takes place only towards the future. If we want to know the decoherence time (let us say, towards the future), we must find the poles (introduced in the previous subsection) in the lower half-plane of the complex energy plane and then consider the pole closer to the real axis²¹: the inverse of its imaginary part is the decoherence time. This means that decoherence towards the future is related to the lower poles (since we are dealing with a process towards classical equilibrium) and, therefore, to the outgoing states. Symmetrically, the (anti)decoherence time towards the past is related with the upper poles (since now we are dealing with a process that takes the system out

²⁰The unstable nature of quantum systems is very well known in quantum computation, where the system's tendency to decohere is the main obstacle to the implementation of information processing hardware that takes advantage of superpositions.

²¹Precisely, the poles of the Von Neuman-Liouville operator, which measure the decaying of the non-diagonal components [46].

of classical equilibrium) and with incoming states. Of course, since the poles are placed in symmetric positions, both decoherence times are equal²². But lower poles are related with $C_+(x)$ and upper poles with $C_-(x)$. Therefore, the time-asymmetry of the universe defines the decoherence direction since it tells us which are the decaying poles (that define the decoherence time) and which are the growing poles (that define the preparation time of the instabilities).

2. Classical statistical mechanics \rightarrow classical mechanics

By means of the Wigner integral, in papers [45], [47] it is shown how the classical statistical mechanical state, obtained by the process explained above, is composed by a set of individual classical states moving along classical trajectories in phase space²³. In fact, the Wigner function $\rho_\omega^W(q, p)$ of the quantum state $(\omega|$ of eq.(7) reads:

$$\rho_\omega^W(q, p) = C \delta(\omega - H^W(q, p)) \delta(a_1 - A_1^W(q, p)) \dots \delta(a_N - A_N^W(q, p)) \quad (9)$$

where C is a normalization constant, $H^W(q, p)$ is the classical Hamiltonian, and $A_1^W(q, p) \dots A_N^W(q, p)$ are the Wigner functions of the remaining commuting observables whose quantum numbers $a_1 \dots a_N$ define the state $(\omega| = (\omega, a_1 \dots a_N|$ ²⁴. Therefore, $\rho_\omega^W(q, p)$ corresponds to a classical trajectory defined by the constants of motion:

$$\omega = H^W(q, p) \quad a_1 = A_1^W(q, p) \quad \dots \quad a_N = A_N^W(q, p) \quad (10)$$

and the Wigner function $\rho_*^W(q, p)$ corresponding to $(\rho_*|$ reads:

$$\rho_*^W(q, p) = \int \rho_\omega \rho_\omega^W(q, p) d\omega \quad (11)$$

This means that $\rho_*^W(q, p)$ is a classical statistical state corresponding to the motions of the classical trajectories $\omega = H^W(q, p)$, $a_1 = A_1^W(q, p)$, ..., $a_N = A_N^W(q, p)$ weighted by the probabilities ρ_ω . To the extent that we can conceive a quantum state as a quantum ensemble²⁵, the quantum ensemble becomes a classical ensemble. Each classical individual state (belonging to the classical ensemble) can be considered as a very small region in phase space if the domain of its density function in phase space is small, even though always satisfying the uncertainty principle $\Delta q \Delta p \gtrsim \hbar$ (we could imagine that it is a coherent state). But if the action S of the local system under consideration is very large ($S \gg \hbar$), this small region becomes almost a geometrical point that follows a classical trajectory for all physical effects related with its size²⁶. As we are now in the classical domain (since in practice $\hbar \rightarrow 0$), we can repeat the argument of Subsection VII.D for the explanation of the process *probability* \rightarrow *fact*, but now starting from a quantum unstable state and going towards its classical limit.

We have not exhausted the list of the arrows of time, but we have shown that the most important ones agree with GAT²⁷.

²²They may be different if a time-reversal non-invariant interaction is present, i.e., weak interaction.

²³Therefore, it corresponds to a Frobenius-Perron evolution [37].

²⁴In equations (3) to (7) we have only written the ω for conciseness.

²⁵A quantum state ρ can always be conceived as a set of individual quantum states since, being selfadjoint, it reads:

$$\rho = \sum_i p_i |i\rangle\langle i| \quad 0 \leq p_i < 1$$

Each $|i\rangle\langle i|$ is an individual state because, considered as a matrix, it has probability 1 (certainty); in other words, it contains maximum information since $Tr(|i\rangle\langle i|^2) = 1$; therefore, if conceived as an ensemble, it would represent a set of copies of the same state $|i\rangle\langle i|$, thus, one $|i\rangle\langle i|$ is sufficient. As a result, ρ can be considered as a set of individual states $|i\rangle\langle i|$ linearly combined in proportions p_i .

²⁶This process can also be studied using Gell'man-Hartle histories, as in Appendix C of paper [45].

²⁷The arrow of time corresponding to the decay of quantum states is treated in paper [3], and the psychological arrow of time in paper [2].

The panorama is not completely closed yet: weak interactions should be included in this scenario²⁸. Most likely they alone will give a complete local explanation of time-asymmetry. However, this fact would not diminish the relevance of the also complete global explanation given by cosmology. From its very beginning, theoretical physics has tried to combine its different chapters in an unified formalism, and it is well known that unifications have always produced great advances in physics. Therefore, our future challenge will be to unify the weak-interactions explanation with the cosmological explanation, instead of abandoning the latter in favor of the former as many local-minded physicists insist.

As it is well known, there is never a last word in physics. Nevertheless, we can provisionally conclude that the global definition of the arrow of time has no serious faults and, therefore, it can be used as a solid basis for studying other problems related with the time-asymmetry of the universe and its sub-systems.

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APPENDIX A: CRITICISM TO SCHULMAN'S ARGUMENT

Schulman [13] exhibits a model in which two weakly coupled systems would maintain opposite running thermodynamic arrows of time. From this model he concludes that regions of opposite running arrows of time at stellar distances from us are possible. This possibility would represent a counter-example to our position: Schulman's model would show that a universe consisting in two weakly coupled sub-universes A and B can have two regional arrows of time pointing to opposite directions.

Even though Schulman's argument sounds convincing at first sight, it becomes implausible when analyzed from a cosmological viewpoint. In Schulman's proposal, the low entropy extremities of the sub-universes A and B are opposed, and both sub-universes evolve towards equilibrium in opposed time directions. Let us consider two cases:

1.- The sub-universe A is bigger than the sub-universe B²⁹ (this situation is not considered by Schulman). If time-orientation is defined by entropy increase, in this case the time-orientation of the whole universe AUB will agree with the time-orientation of A, and B will go from equilibrium to non-equilibrium. Nevertheless, the behavior of B is neither strange nor unnatural: since there is a flux of energy which, according to the time-orientation adopted, must be considered as a flux from A to B, then we can consider that it is such energy what takes the sub-universe B out of equilibrium. In other words, the decreasing entropy of the *open* sub-universe B has the same explanation as the decreasing entropy in the usual open systems that we find in our everyday life.

2.- The sub-universe A is equal to B (the situation studied by Schulman, where A and B are identical). In this case, the universe AUB is perfectly time-symmetric. But, as the theorem of Section V has shown, time-symmetry has vanishing measure: it requires an overwhelmingly improbable fine-tuning of all the state variables of the universe.

But even in the time-symmetric case, it is not admissible to suppose that the sub-universes A and B have opposite time-orientations. When considered as a cosmological model, Schulman's simple model describes a time-orientable universe. In a time-orientable manifold, continuous time-like transport has conceptual priority over any method of defining time-orientation. In other words, Schulman's universe should have a light-cone structure such that, if we continuously transport a future pointing vector from the point $x \in A$ along some curve to the point $y \in B$, the transported vector will fall into the future lobe $C_+(y)$. This means that A's future cannot be different than B's future: there is an only future for the whole manifold, defined by its light-cone structure.

However, who prefers to insist on the attempt to use entropy increase for defining time-orientation could appeal to the following strategy: to define the future direction of time as the direction of the entropy increase, for instance, in

²⁸ K_0 and \overline{K}_0 are not twins; therefore, they introduce a completely different arrow of time, the weak interaction arrow [15], [39].

²⁹To fix the ideas, we can say "externally bigger". The cases "more or less bigger" or "almost symmetric" can be included in the coarse-grained version of the corkscrew theorem, since in these cases there is only a small difference with a time-symmetric model. Then, these solutions have small measure.

the sub-universe A, and then to establish the time-orientation in the sub-universe B by means of continuous time-like transport. But who adopts this strategy is committed to explain why future is the direction of entropy increase in one region of the universe but not in the other: why the entropy definition works in one region of the universe but not in all of them. These considerations lead us to our starting point: the problem of the arrow of time should be addressed from a global perspective, taking into account the geometrical properties of space-time.

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